arXiv:astro-ph/9610262v1 31 Oct 1996

The CATS database to operate with astrophysical catalogs

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Abstract.   A public database of astrophysical (radio and other) catalogs (CATS), has been created at Special Astrophysical Observatory (SAO). It allows to execute a number of operations in batch or interactive mode, e.g. to obtain a list and parameters of catalogs, to extract objects from one or several catalogs by various selection criteria, perform cross-identification of different catalogs, or construct radio spectra of selected sources. Access to CATS is provided in both dialog mode (non-graphics), and graphics mode (hypertext, via Tcl/Tk or possibly JAVA in future). The result of CATS operation can be sent to the user in tabular and graphical formats.

1.   Introduction

Different attempts have been made for combining many astronomical catalogs in unified databases NED, SIMBAD, ESIS, ADS, etc. (see reviews by Andernach et al. 1994; Andernach 1995). Important shortcomings of these databases are (a) the incompleteness of the information stored compared to that offered in the original publication and (b) the necessity to copy the whole catalog (if available at all) for dedicated work with it. We propose a new solution to this problem with a "CATalogs supporting System" (CATS) at Special Astrophysical Observatory (SAO). It allows to operate with catalogs coded in plain ASCII, to cross-identify different radio catalogs, to calculate spectral indices, to construct and fit spectra. This database is now running on the server *ratan.sao.ru* of SAO of Russian Academy of Sciences (RAS), and constitutes part of the data bank of SAO RAS (Kononov, 1995).

2.   Realization of the database

The present active database of catalogs of version 1.0 (Fig.1) is a unification of catalogs, descriptions of catalogs and programs operating with them under the freely distributed UNIX version OS Linux at the RATAN–600 server with
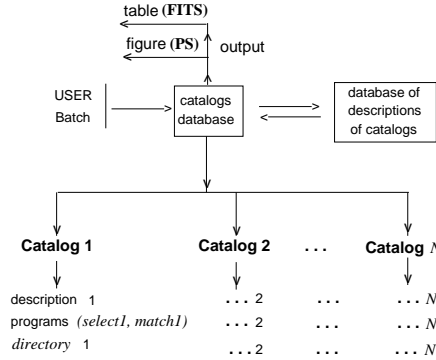
table (**FITS**)

figure (**PS**) ↑ output

USER Batch → catalogs database ⇄ database of descriptions of catalogs

Catalog 1   **Catalog 2**   . . .   **Catalog** N

description 1   . . . 2   . . .   . . . N
programs (*select1, match1*)   . . . 2   . . .   . . . N
directory 1   . . . 2   . . .   . . . N

Figure 1.   The scheme of the CATS database.

a Pentium processor (Verkhodanov & Trushkin 1995). The programs are coded in "C" and are freely shared except for commercial goals.

New catalogs may be added to CATS in conformity with the following rules: 1) every new catalog of objects has to be contained in the UNIX directory having the same name as the catalog of objects; 2) the description of the catalog must also be in that directory; 3) the programs for local operations with the catalog of objects are also in the same directory; 4) brief characteristics and names of the programs and file with the description of the catalog are stored in the file `cats_descr`. The described manner of catalog storage eases the database development, its expansion with new data and the fine-tuning of the supporting programs.

Virtually all catalogs have a different format and list different observables. It will be a major challenge to provide uniform access to such a heterogeneous collection of data sets based on different methods, using different notations and units (in the absence of a "standard" to create catalogs). Except for parameter-dependent derived quantities, we intend to use all different fields (i.e. columns of data tables) as they were published.

3.   Capabilities and Access

CATS is capable to accomplish the following tasks: 1) Provide a short description and characteristics of each catalog; print the full list of catalogs relevant for a given sky area. 2) Select objects from one or several catalogs matching user-specified criteria, like equatorial and galactic coordinates, flux densities and spectral indices, observing frequency, names of catalogs unified in mixed catalogs as Dixon's Master Source List, and object type (if provided by the catalog). 3) Cross–identification of different catalogs and selection after calculation of spectral indices. 4) Drawing radio spectra of selected sources in PostScript.

The result of the object selection can be sent to the standard output or saved in the following formats: 1) The original format of the input catalog. 2) A standard format, common for all catalogs and used further for unification and operation with spectra or other parameters. The standard FITS TABLE header
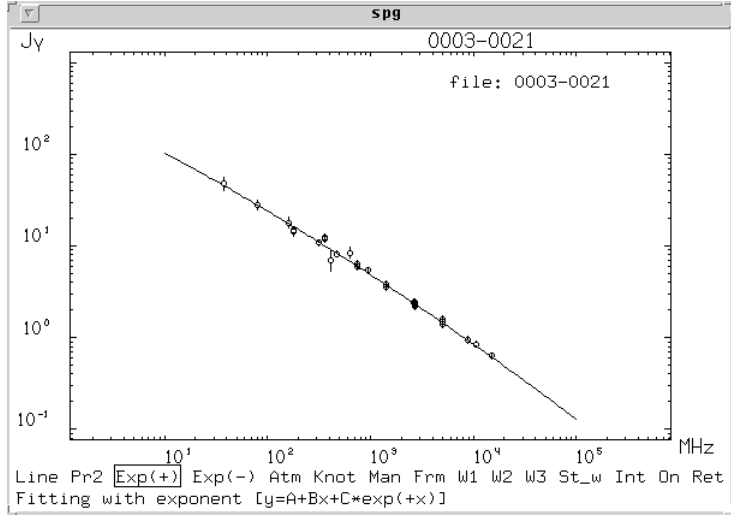
Figure 2.   The *spg* screen with the menu and the radio spectrum.

describing the various data fields of the table may be recorded with the resulting file. 3) X-window codes or tape archives (TAR format) of compressed PostScript files for graphical spectra of radio sources. The result of the CATS operation is an 'ASCII'-file sorted according to different object characteristics. It can be used for subsequent investigation of the radio source spectra or statistical source properties in the RATAN–600 data processing system (Verkhodanov, 1997).

On-line access to CATS will be provided in several modes: 1) Dialog mode (non-graphics) is the only mode established until now. Several UNIX *shell* scripts (Verkhodanov & Trushkin 1995) permit to run the database-supporting programs via TCP/IP and NFS protocols in the local computer net of SAO. 2) Access via TCL/Tk scripts on the basis of *shell*. This mode will also allow to operate with the figures or profiles of radio spectra or statistical distributions by different parameters of selected samples. 3) Hypertext access will eventually allow remote Internet users to operate with CATS via hypertext transfer protocol (HTML). It will allow to execute all described operations in graphic mode and probably take advantage of the JAVA language.

Presently we are working on a further type of access to CATS by e-mail request enabling the user to send a message with his/her requests. The latter will be read automatically and sent for execution to the CATS scripts. The result will be sent automatically to the user via e-mail. Very bulky results will be placed in a public FTP area and the user will be informed about the FTP address of the file(s). The e-mail messages may have several formats describing the search window, the coordinates of the center, the epoch, the type of output format, and the type of catalogs in which to search. CATS also permits to copy entire catalogs (via FTP) to the user's local computer and to operate with them "at home".

CATS operates with continuum spectra of radio sources recorded as FITS tables. The graphics program *spg* (SPectra Graphics) allows to fit a spectrum

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with a standard set of curves: 1) $y = A + Bx$, 2) $y = A + Bx + Cx^2$, 3) $y = A + Bx + C \times exp(x)$, 4) $y = A + Bx + C \times exp(-x)$, where $x = log\ \nu$, $y = log\ S$, ν is the frequency (MHz), S is the flux density (Jy).

The menu of the *spg* program (Fig. 2) allows to choose among these curves either automatically (by a least-squares fit), or by manual selection of the fitting function, or by manual fitting using the mouse (when the curve follows a cursor). Data points may be weighted in different manners: setting equal weight, setting weights by flux density errors or filling a form with a table of frequencies, flux densities and weights.

4.   Conclusions

The development of CATS will provide a simple and convenient access to astrophysical information and accelerate the process of obtaining characteristics of celestial objects. Operation with the database will permit astronomers to search for peculiar objects and study physical processes in sources of cosmic radiation.

CATS allows not only to get accurate positions of radio sources and to study radio spectra, but also to derive different statistical properties of object samples. Trushkin & Verkhodanov (1995) recently demonstrated such possibilies in a cross–identification of two large catalogs in two different frequency ranges: the IRAS Point Source Catalog in the infrared and the UTRAO survey at 365 MHz. Two of us (H.A. and O.V.) are working on the cross–identification and eventual optical identification of UTR survey sources detected at 12–25 MHz. We plan to solve the problem of very large error boxes by a stepwise cross-ID progressing from low frequencies and angular resolution to higher ones until the error box size permits the optical ID.

CATS is being expanded continuously. Presently it comprises over 50 catalogs including all the RATAN–600 catalogs and occupies $\sim$250 Mb. The database system could be an essential part of a bigger project of the first publicly accessible database of radio sources, proposed by Andernach et al. (1997).

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